

PRELIMINAR ORBITAL ELEMENTS OF THE ECLIPSING BINARY SYSTEM HD 217 796 = NN CEPHEI

The eclipsing character of the binary system HD 217 796 was first established by GEOS observers in 1976 (see GEOS NC 142 and IBVS 1231).

From the published light-curves (NC 142, NC 159) and from the additional one shown in figure 1 (obtained using my own observations), I have computed the rectified light-curve of the figure 2, by an harmonic analysis: The evidence of the arch between eclipses is poor, that indicates weakly ellipsoidal stars.

From the rectified light-curve, I have derived the geometric and the photometric orbital elements according to the TSESEVICH method for a limb-darkening coefficient $x = 0.5$ (and for spectral type A5).

The orbital elements computed for a partial eclipse with the primary minimum corresponding to an occultation (large star in front) are showed in table 1. A representation of the binary system NN Cep is given in figure 3.

Conclusion: NN Cep is probably an EB - type system, but I do not exclude the EA - type because the ellipticity of the stars are very small and the values of the radii of the two bodies could be representative of a detached binary system of the algol - type. New and more precise observations are required.

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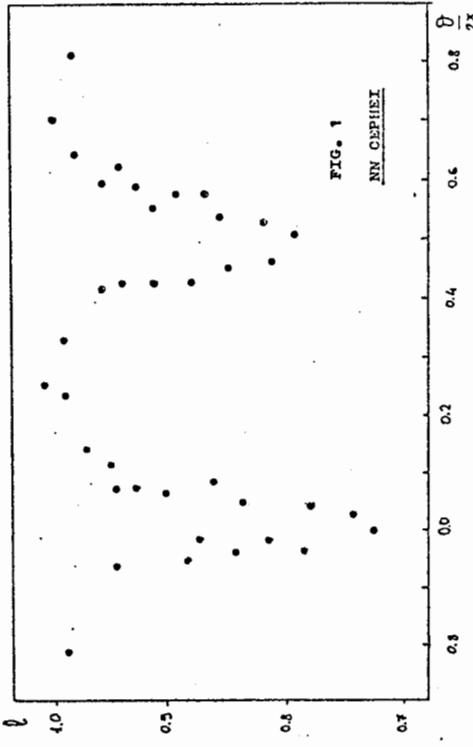
GEOMETRIC ORBITAL ELEMENTS

orbital eccentricity	$e = 0.026$
longitude of periastron	$\omega = 76^\circ 04'3$
orbital inclination	$i = 80^\circ 51'6$
ratio of radii	$k = 0.575$
longer radius of the large star	$r_1 = 0.293$
longer radius of the small star	$r_2 = 0.168$
shorter radius of the large star	$b_1 = 0.286$
shorter radius of the small star	$b_2 = 0.164$
ellipticity	$\epsilon = 0.976$
occulted area	$a_{\text{oc}}^{\text{oc}} = 0.903$
transited area	$a_{\text{oc}}^{\text{tr}} = 0.859$

PHOTOMETRIC ORBITAL ELEMENTS

light of large star	$L = 0.715$
light of small star	$L = 0.285$
ratio of surface brightness	$J = 1.205$
limb darkening coefficient	$x = 0.5$ (assumed)

Table 1 . Geometric and photometric orbital elements of NN Cep.

FIG. 1
NN CEPHEIFig. 1 - Composite by GAS ($\tau = 155$) 1977 .

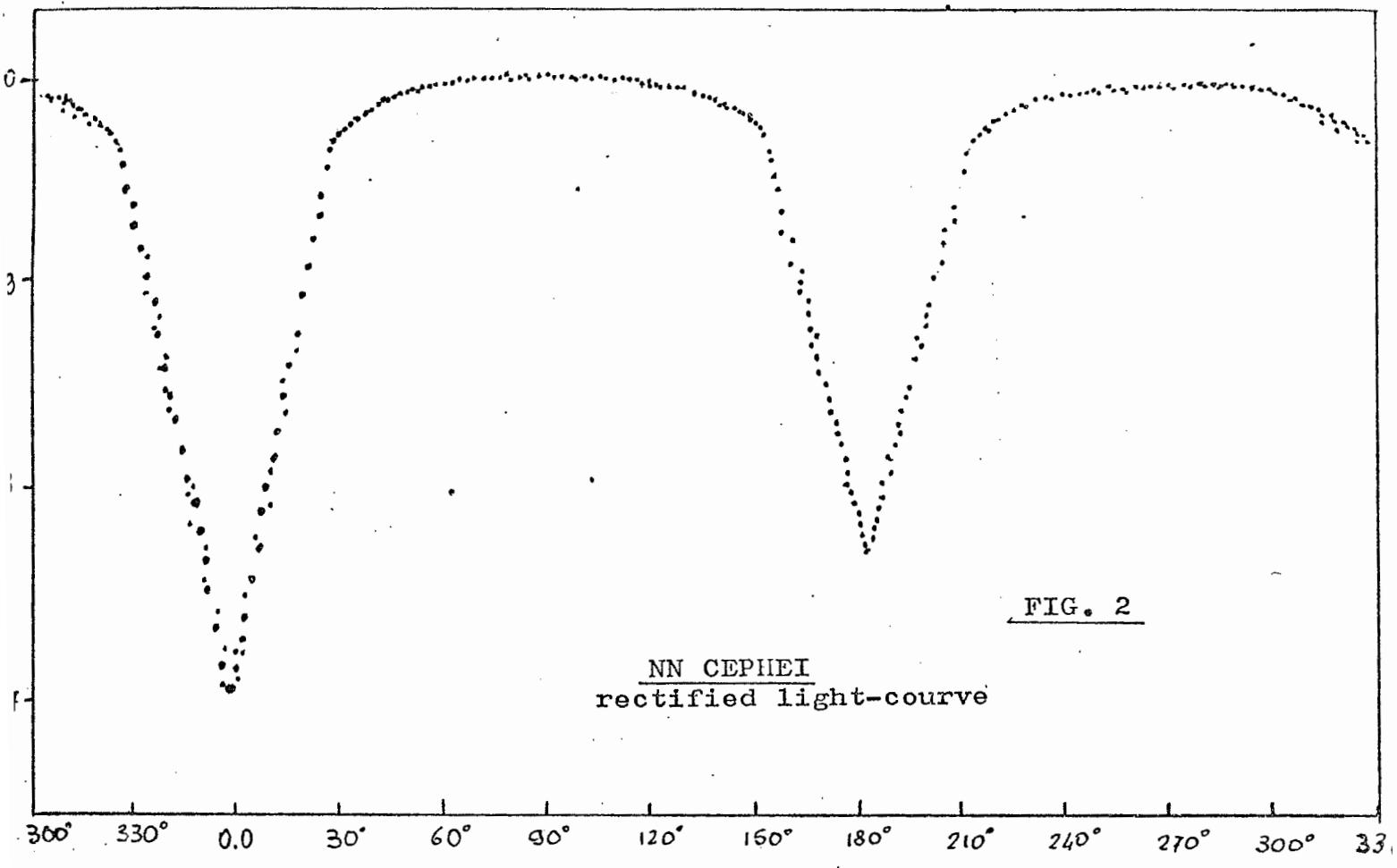


FIG. 2

NN CEPHEI
rectified light-curve

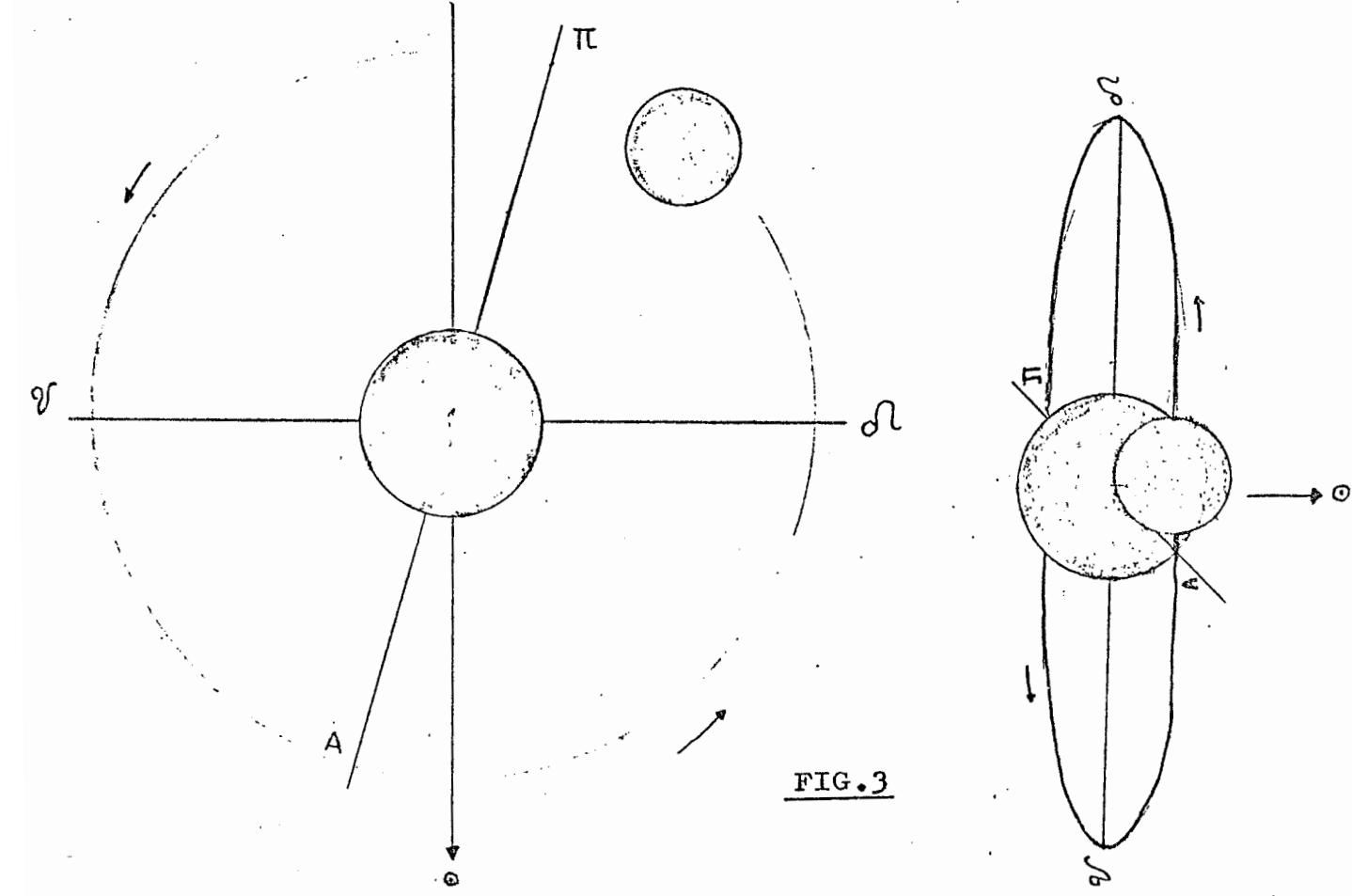


FIG. 3

Graphic representation of the
binary system IID 217796